



Measuring the age of planetary surfaces using crater statistics: Guide to the *Craterstats* software

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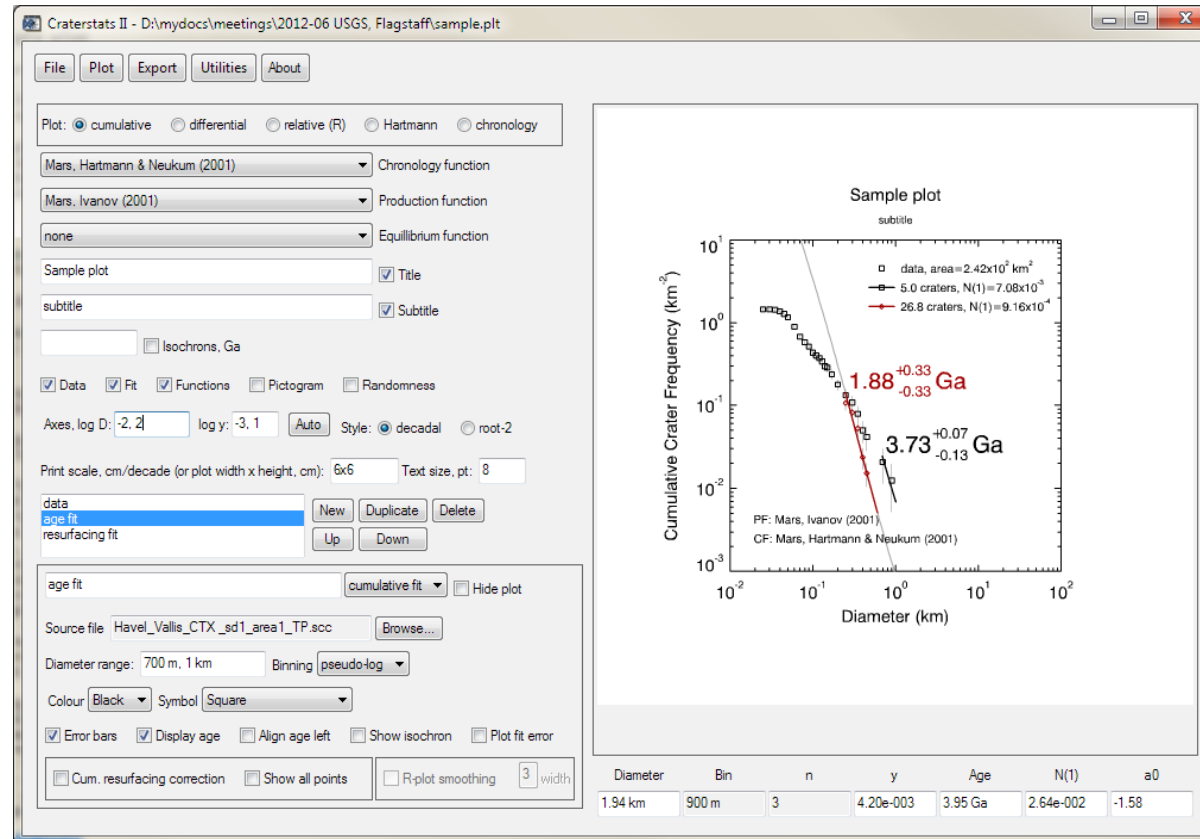
Freie Universitaet Berlin

Switch view

Global settings

Plot list

Plot settings

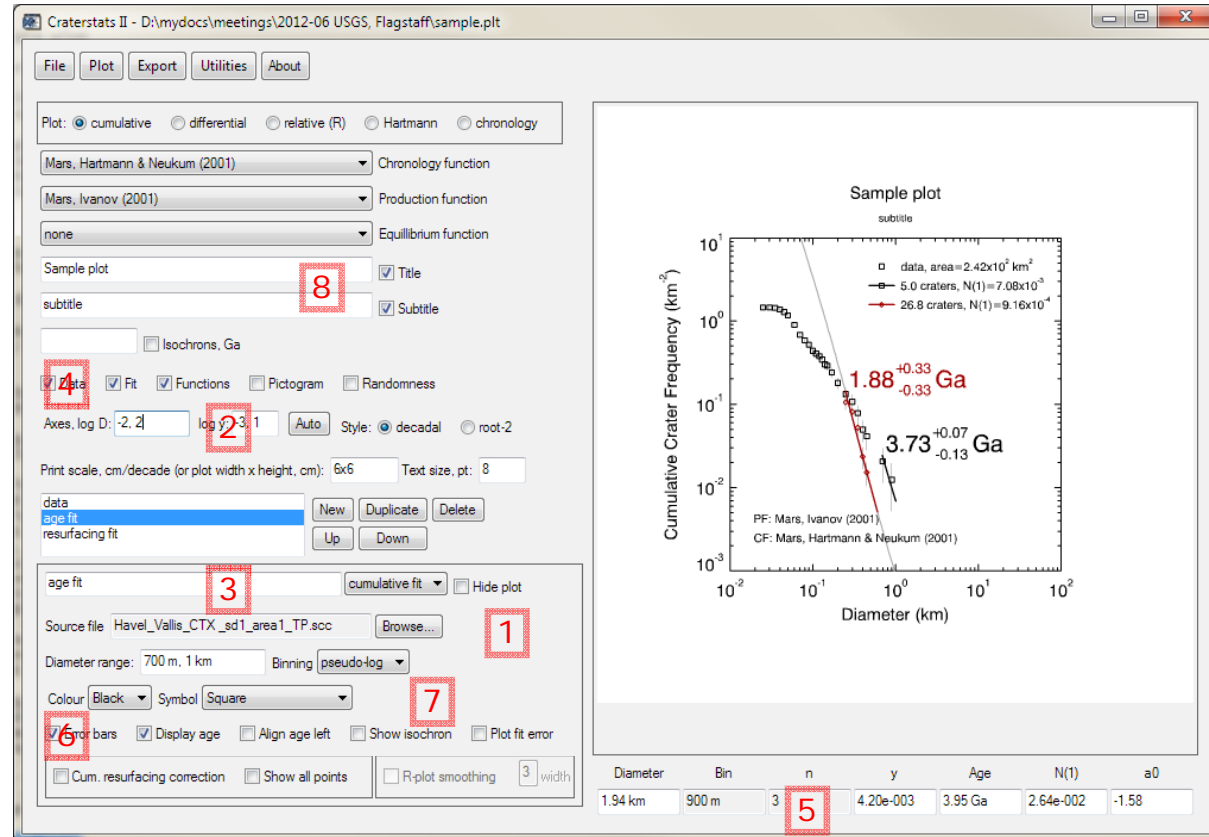


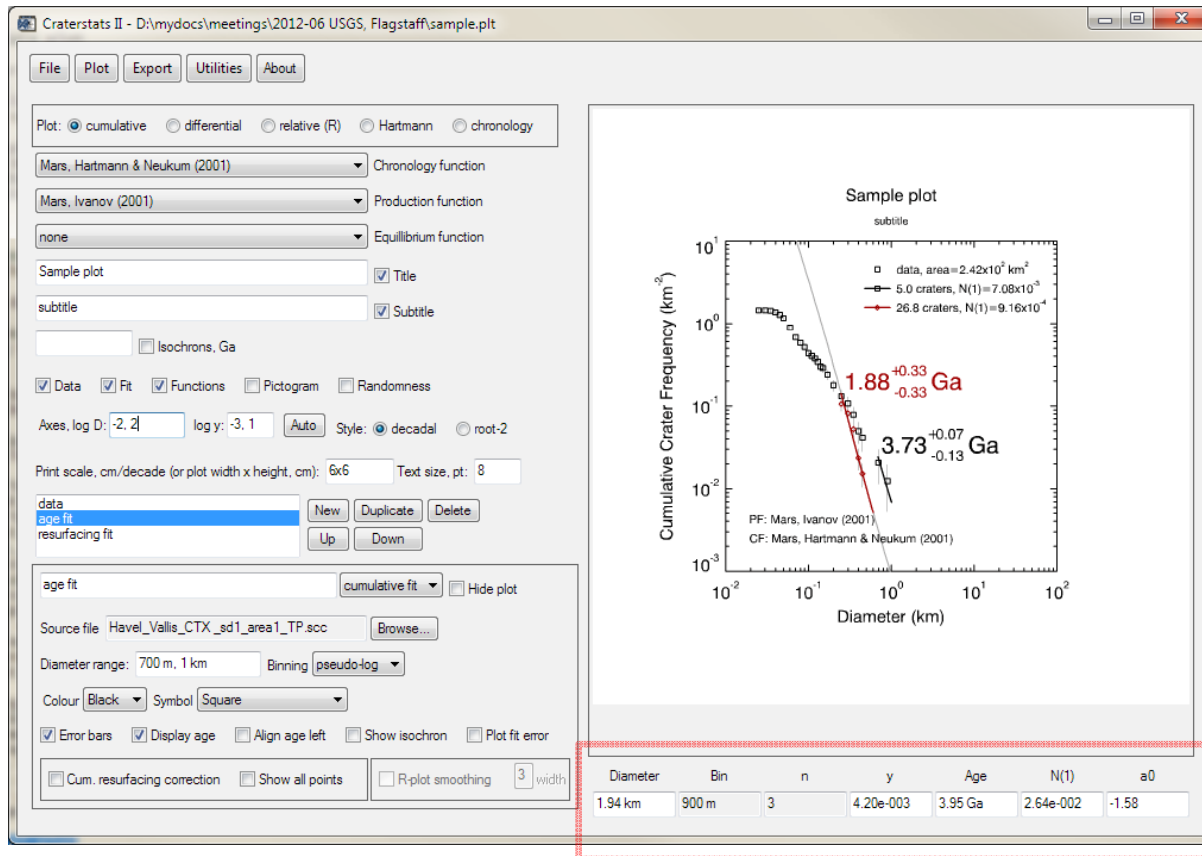
Plot window

Info bar

Create a data plot

1. open source data (e.g. sample.diam)
2. set axes (manual/auto)
3. change plot name
4. display data legend
5. note displayed values in info bar on moving cursor
6. note error bars (on/off)
7. plot symbols/colours
8. title/subtitle





y – crater density in no./km² in cumulative view

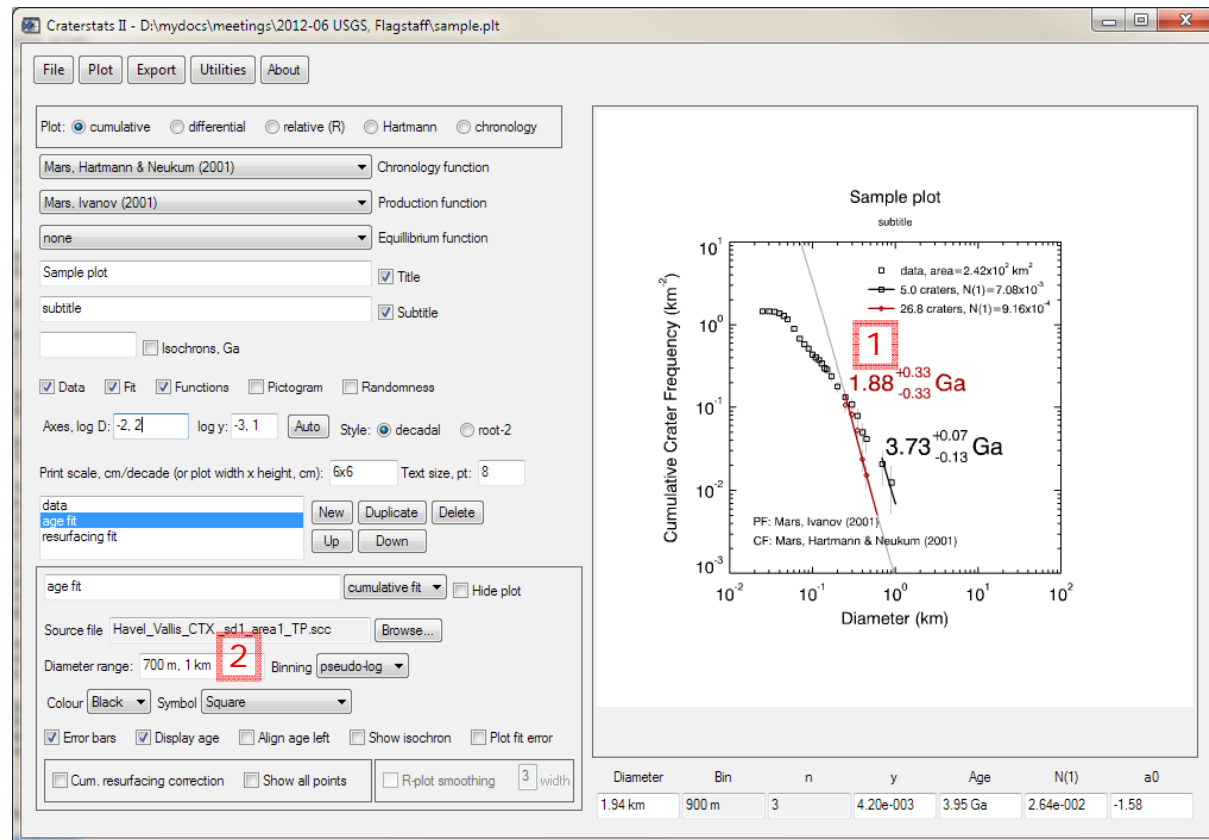
How old is a surface with 96 >5 km craters per million sq. km?

What's the density of >100m craters on a 1 Ma surface?

Enter known values: corresponding ones are computed for current production/chronology functions

Selecting data range

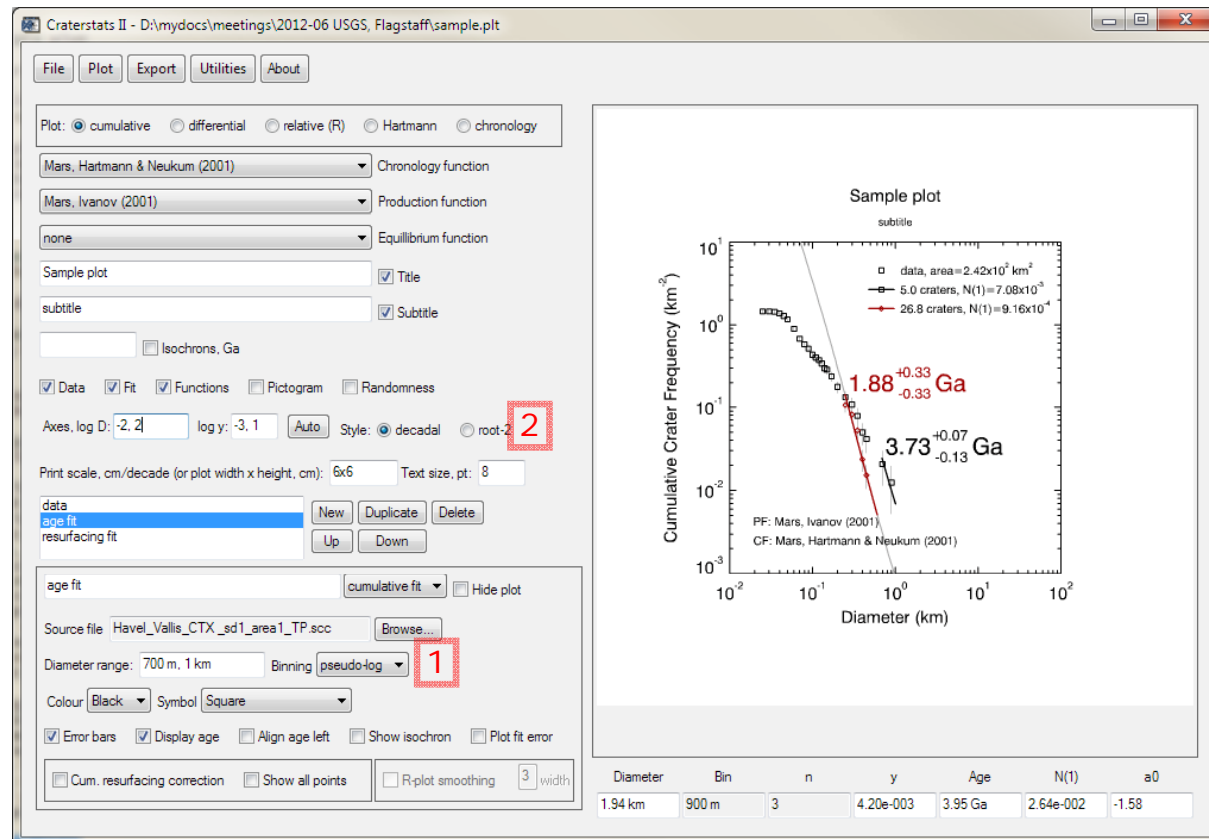
1. Click and drag in plot window, or...
2. Enter range manually



1. Binning

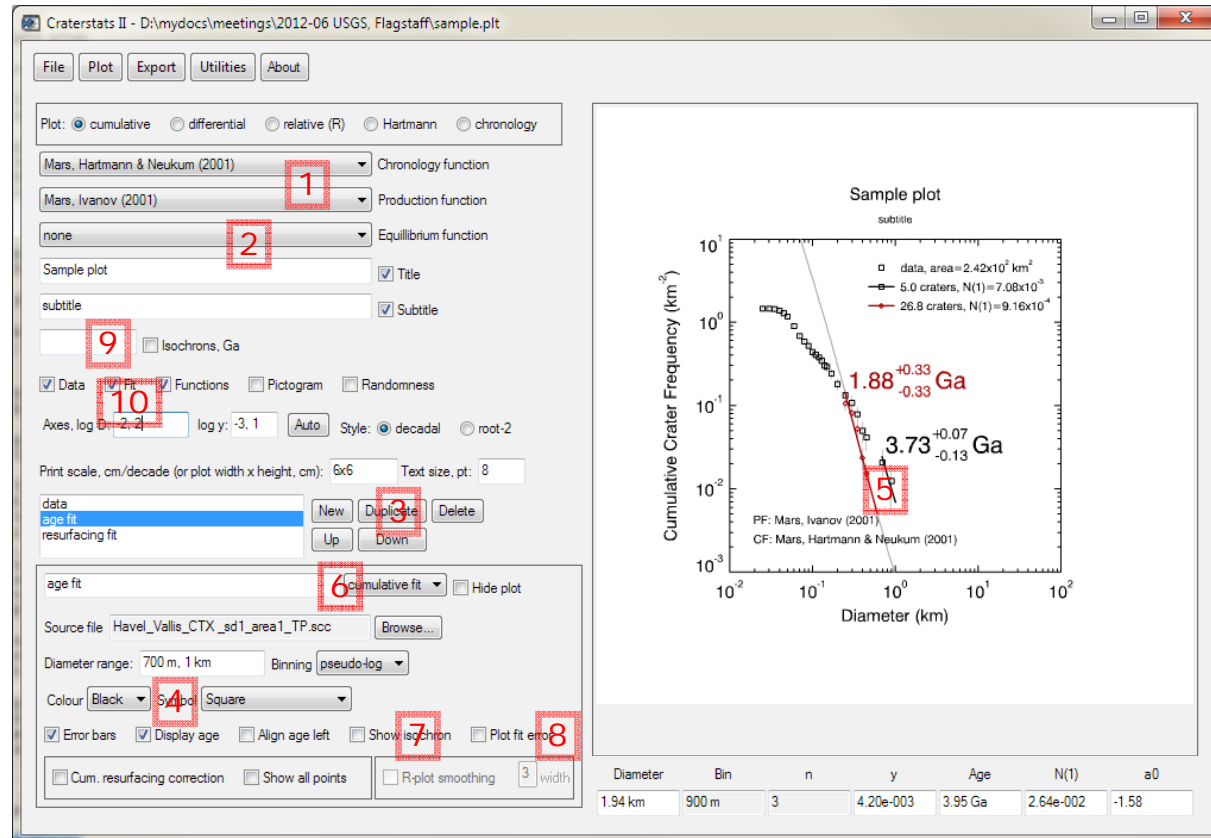
- Pseudo-log - Neukum's technique, 18 bins/decade at rounded intervals, roughly evenly spaced
- Root-2 (Hartmann and others' technique: boundaries at 1, 1.4, 2, 2.8, 4...)
- None – cumulative plot can be made without any data binning
- Others - e.g. 20/decade - like Neukum's, but regularly spaced; 4th root-2 - twice as finely spaced as root-2

2. Axis style: decadal vs root-2 - no influence on data handling: just display



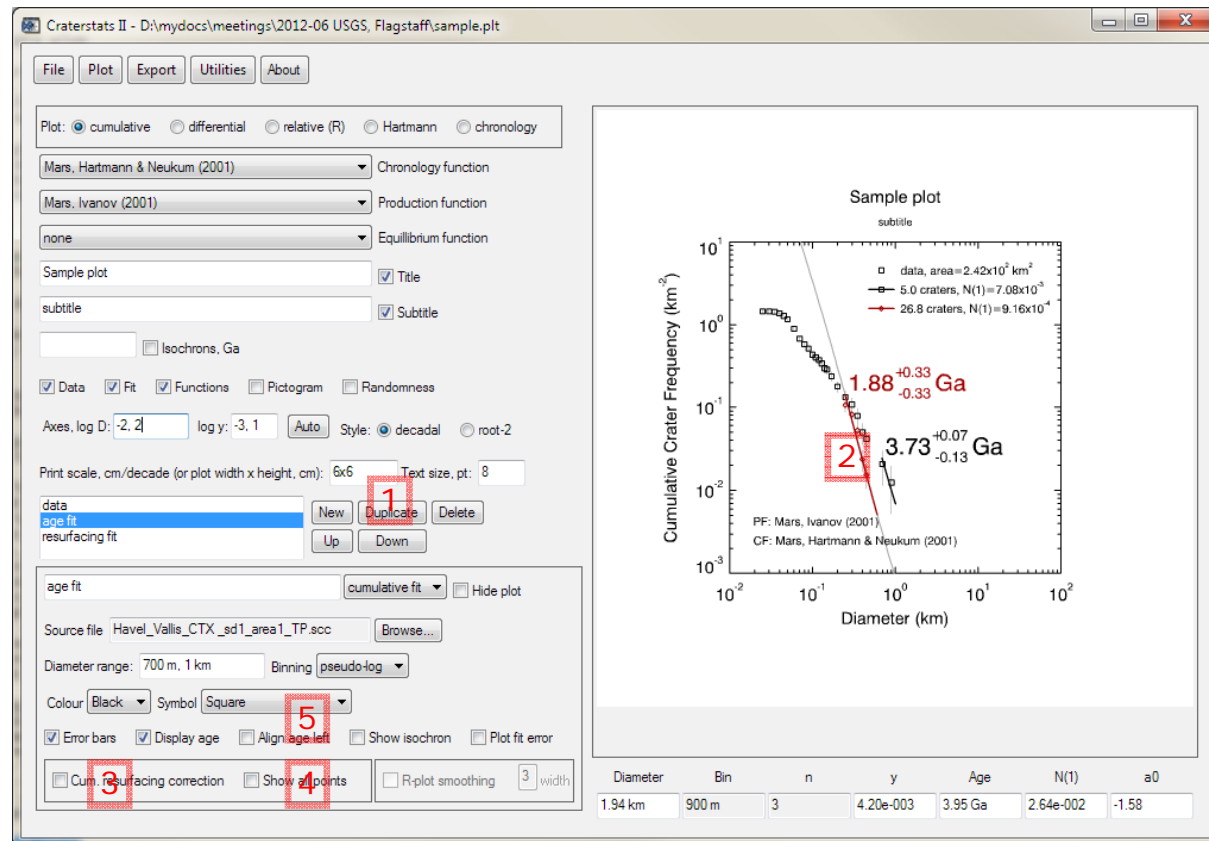
Isochron fitting

1. Select production and chronology functions
2. Show equilibrium function
3. Duplicate existing data plot
4. Change to red triangles
5. Select diameter range (click-drag in plot)
6. Switch plot type to cumulative fit
7. Show complete isochron
8. Plot fit error (on/off)
9. Add user-selected isochron
10. Show fit legend (on/off)



Making a resurfacing fit

1. duplicate existing fit (over largest craters)
2. select new range
3. **apply resurfacing correction:** accounts for the excess large craters present in the cumulative function (Michael & Neukum, 2010). Normally reduces the model age with respect to an uncorrected fit
4. show all corrected points (on/off): we normally show only the corrected points of the fit range, which lie below those of the uncorrected data
5. align age left (on/off)

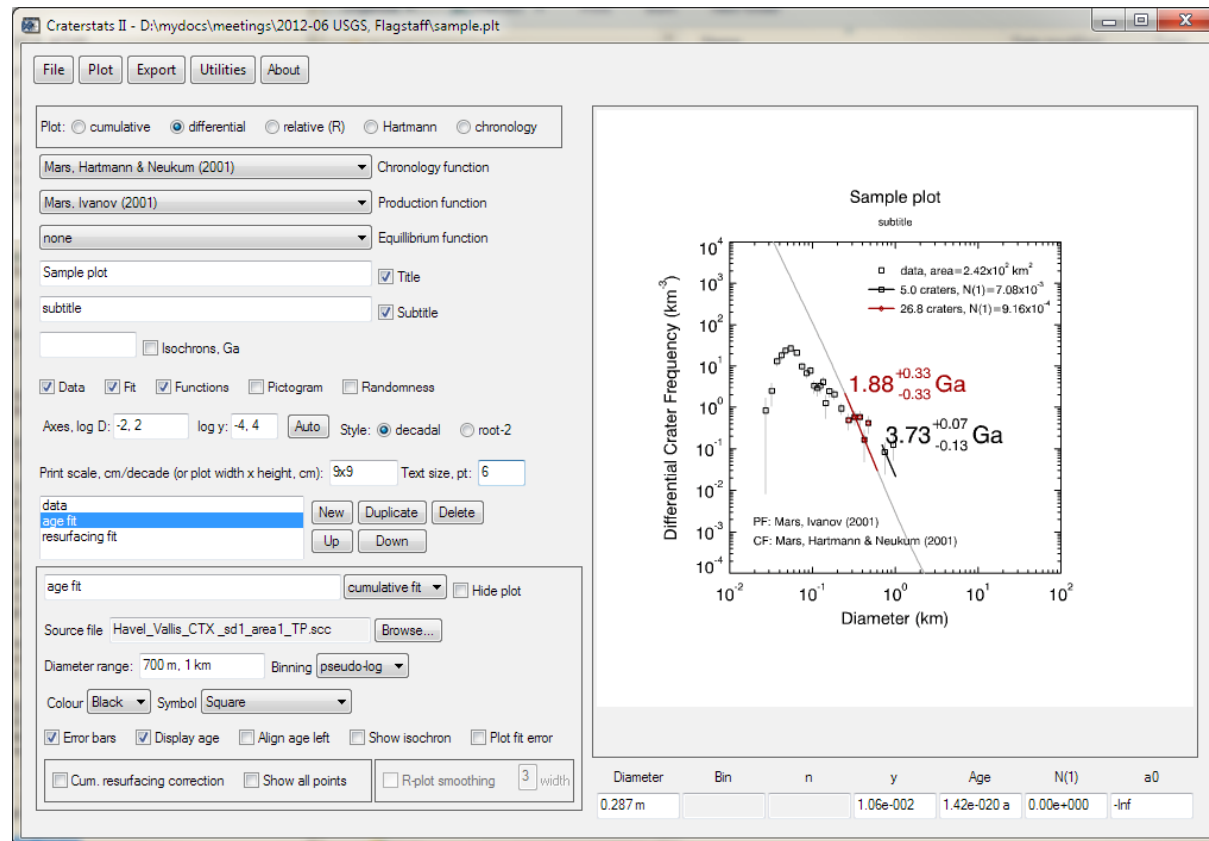


Differential plot

Plots craters per bin
bin width

the bin width term makes the plot essentially independent of the type of binning used. Similar to a Hartmann plot, which plots specifically the number of craters per root-2 bin

very useful for considering diameter range attributable to a resurfacing event

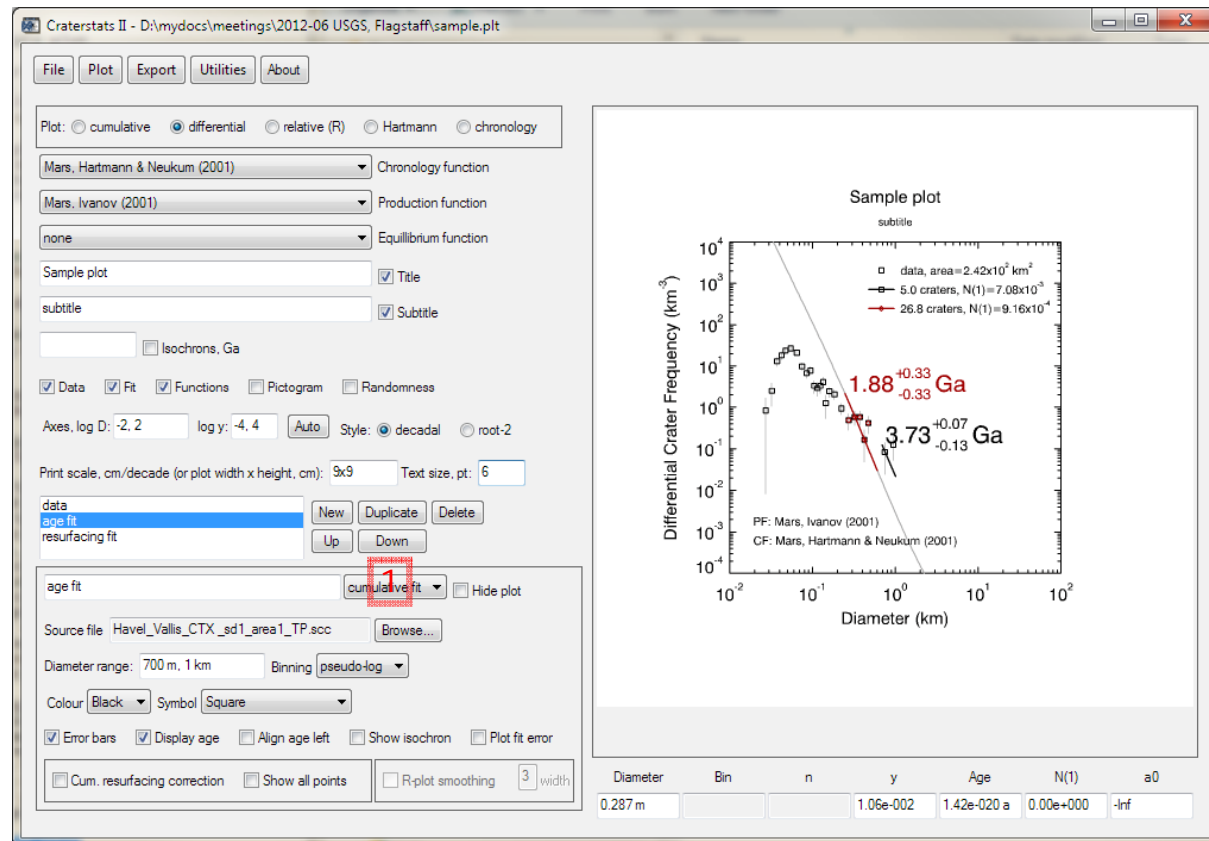


Differential fit

1. may use differential form of production function (Michael & Neukum, 2010) to fit the data

for "ideal" dataset, should obtain same result as cumulative fit

practically, there are small differences, due to different weighting of data points (in a cumulative plot, each crater is represented in more than one data point)

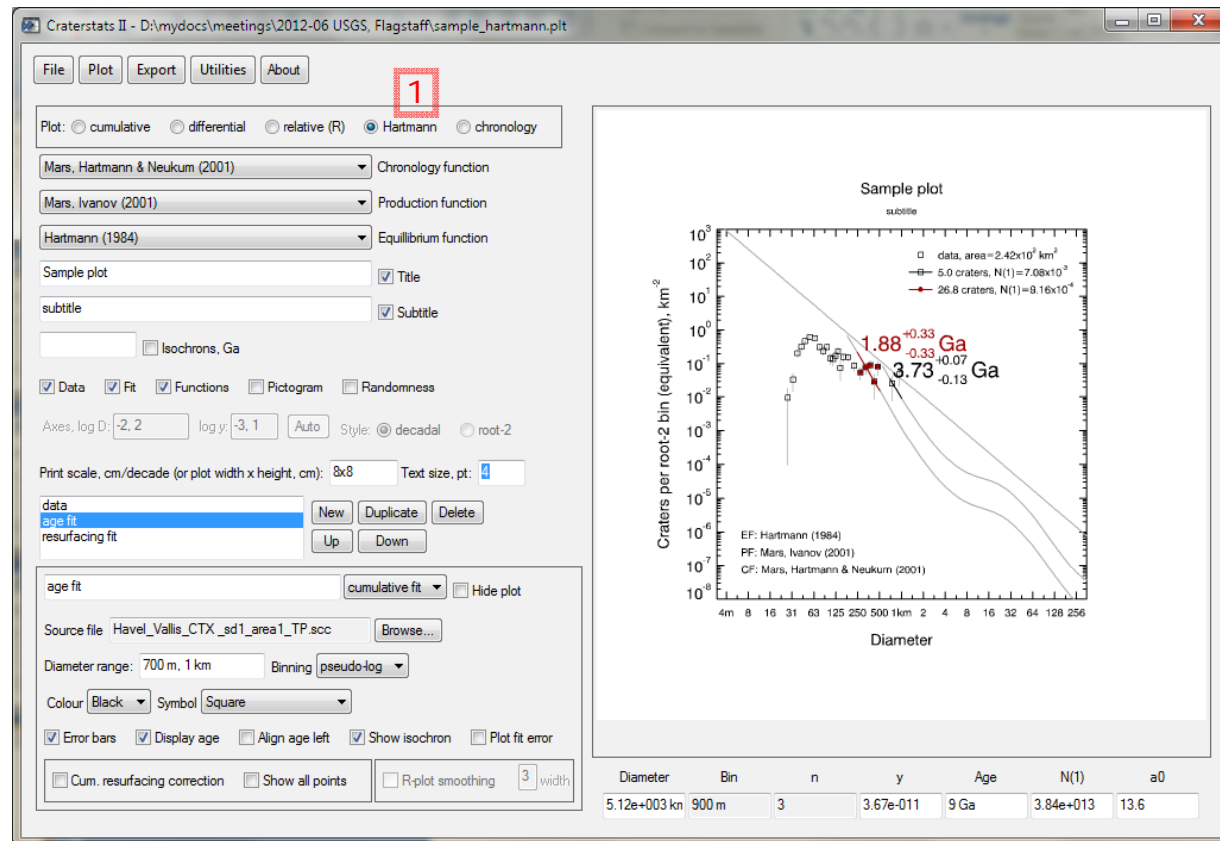


Hartmann plot

1. Hartmann plot is a special case of the differential: here, craterstats plots craters per root-2 bin "equivalent": non-root-2 binnings are normalised to be directly comparable with a Hartmann plot (Michael & Neukum, 2010)

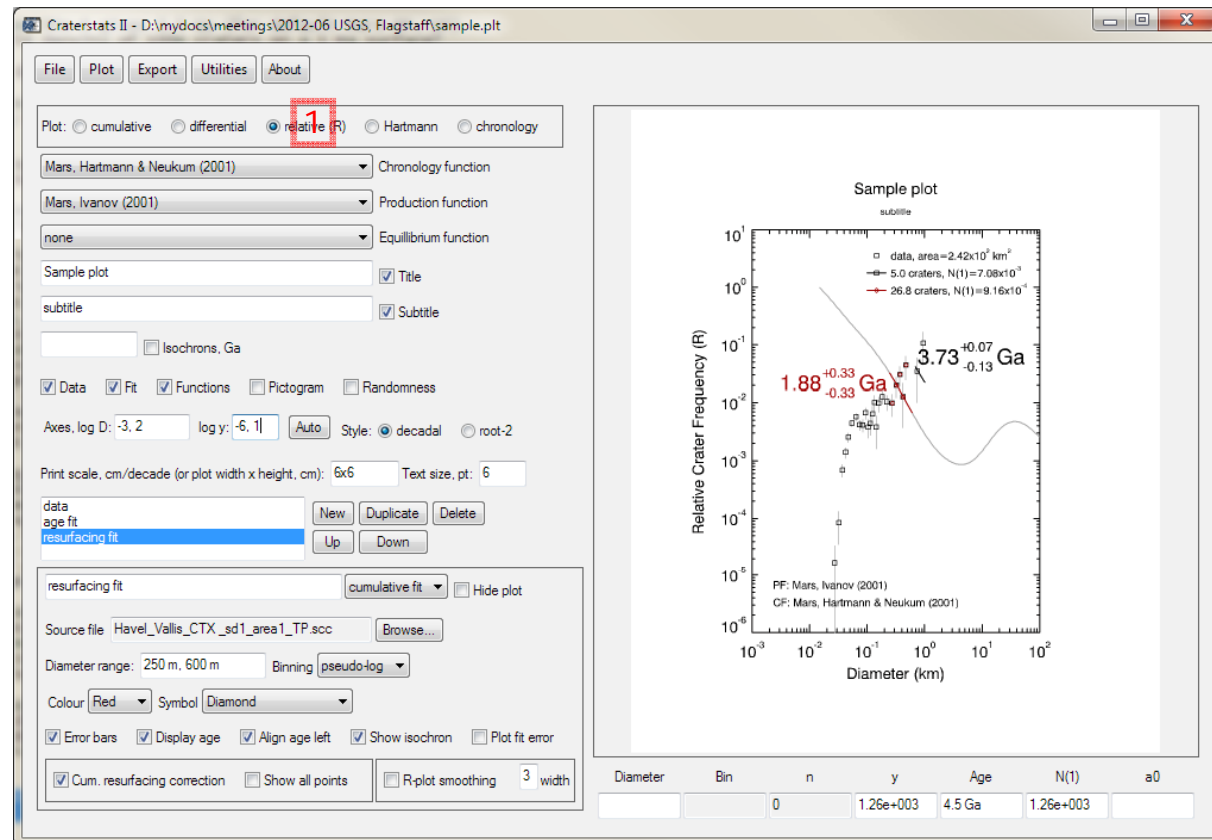
Software handles Hartmann piecewise (and tabular) production functions, but no chronology function at present for Hartmann system (to be added shortly)

Can plot Neukum-style PF/CF results on a Hartmann plot and v.v.



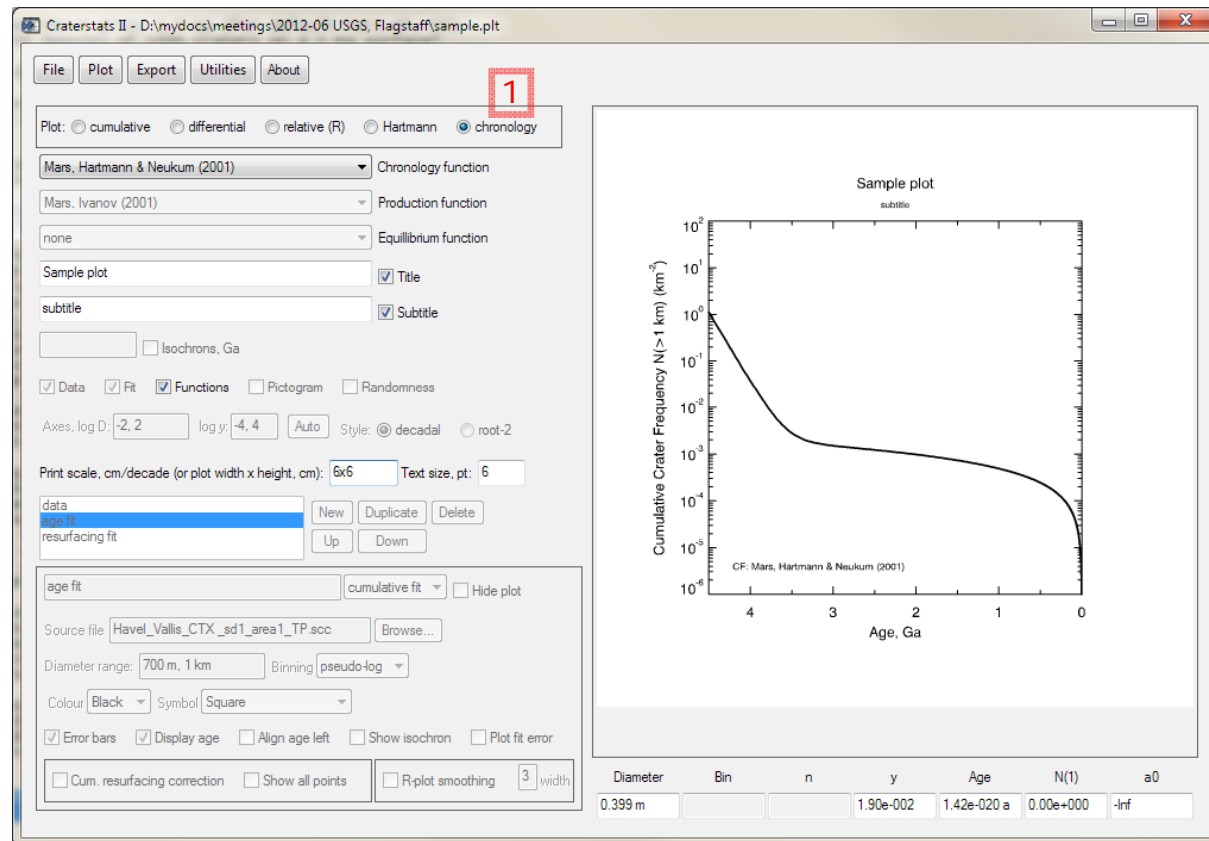
1. Plots the slope relative to a differential -3 slope

Favoured for comparing size-frequency characteristics of crater or impactor populations



Chronology plot

1. not a data view like the others, but just a plot of the current chronology function



File

- Save: save the current collection of plots in a form which can be opened and edited (.plt file)
- Open: open such a file
- Close
- Exit

Plot

- New/duplicate/delete - repetition of the plot list buttons

Export

- Image (either .png or .eps of current view)
- Summary file - text file with key data from current plots and fits. Can be read, e.g. into Excel table
- .stat table (binned data table)

Utilities

- sum .stat files (merge pre-binned data)
- merge .diam files (merge crater lists - preferred)
- randomness analysis - run randomness analysis program

About - program info

In the program folder are several text files which may be modified to alter the default settings:

craterstats2.ini

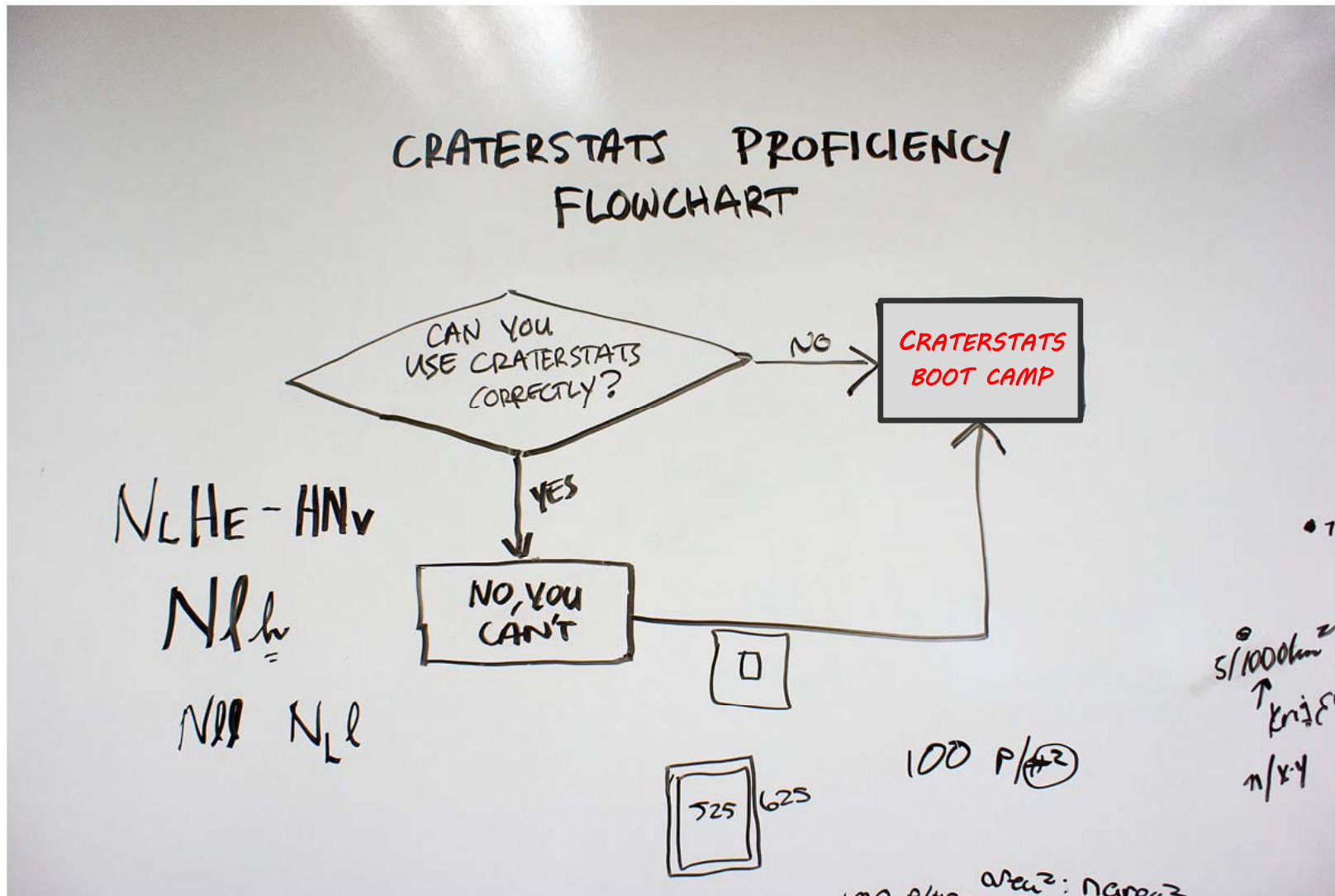
a few control settings for the program, e.g. window size, starting folder for file dialogs

default.plt

plot settings: edit this to change the default axis range, view, production and chronology functions, isochrons etc.

functions_user.txt

here you can define new production functions (polynomial or piecewise), chronology functions or equilibrium functions. The format should be modelled after those present in functions.txt



- Start

- Download from:
 - <http://hrscview.fu-berlin.de/software.html>
- Install IDL virtual machine
- On Windows, double-click craterstats.sav; on linux: type
idl -vm=craterstats.sav

- Make cumulative histogram plot

- Use pre-binned count (Echus Chasma, Mary Chapman):
 - [sample/ECHUS_CHAP_AR1_A.stat](#)
- Select the source file for the count using the 'Browse...' button
- Set the axis ranges

- Make first fit

- Duplicate first plot (File-duplicate), and change plot-type to 'fit'
- Choose production function
- Select fit range
- Options: age, error bars, isochron

- Make 2nd fit (younger part)

- Duplicate previous plot
- Modify fit range
- Hide plot

- Make resurfacing correction

- Duplicate first plot
- Change colour/symbol
- Modify diameter range
- Enter minimum fitting diameter for correction
- Check 'resurfacing correction' box

- Make fit to corrected points

- Duplicate previous plot
- Change plot type to 'fit'
- Set minimum diameter equal to the 'minimum fitting diameter'
- Display age/isochron
- Option: align left

- Export graphic

- Export image as .png file, or as postscript
Save composite plot (File-Save...)
- Examine .plt and .txt output files (both ASCII)

- Now try with your own crater count